

# Properties of symmetric and asymmetric nuclear matter within DBHF theory

Christian Fuchs

Institut für Theoretische Physik

EBERHARD KARLS  
UNIVERSITÄT  
TÜBINGEN



ANL EOS workshop

August 2008

# Outline

Introduction

Relativistic dynamics

Dirac-Brueckner-Hartree-Fock

Nuclear bulk properties

Proton-neutron mass splitting

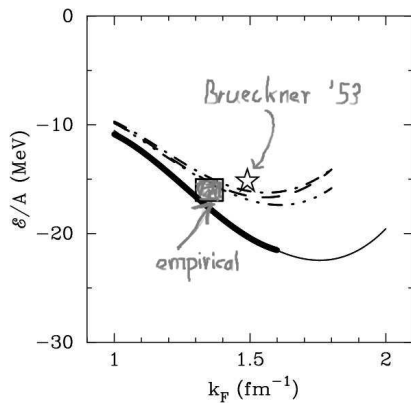
DBHF versus BHF

Chiral condensate

Heavy ion reactions

Summary

## Quest for the EOS

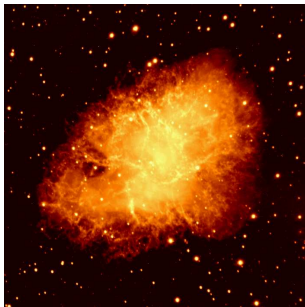


- ▶ EOS at high density ?
- ▶ EOS at extreme isospin ?

## Creation of super-dense matter

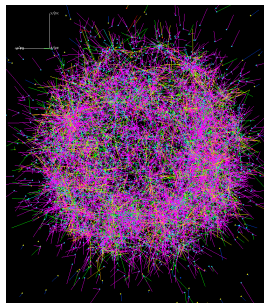
One has to compress by

Gravitation



$$\rho \simeq 10\rho_0$$

Kinetic energy

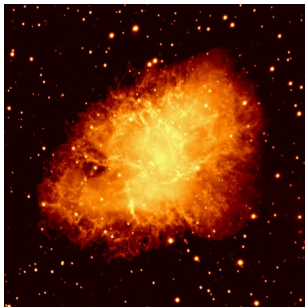


$$\text{SIS: } \rho \simeq 3\rho_0, \text{ FAIR: } \rho \simeq 8\rho_0$$

## Creation of super-dense matter

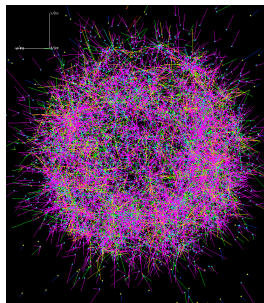
One has to compress by

Gravitation



$$\rho \simeq 10\rho_0$$

Kinetic energy

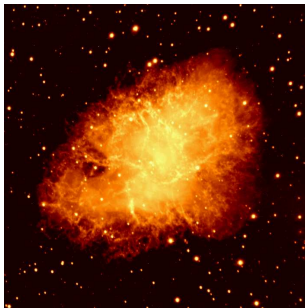


$$\text{SIS: } \rho \simeq 3\rho_0, \text{ FAIR: } \rho \simeq 8\rho_0$$

## Creation of super-dense matter

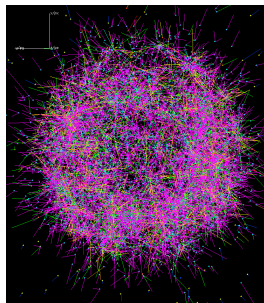
One has to compress by

Gravitation



$$\rho \simeq 10\rho_0$$

Kinetic energy



$$\text{SIS: } \rho \simeq 3\rho_0, \text{ FAIR: } \rho \simeq 8\rho_0$$

## Overview models

- ▶ **Ab initio approaches**  
Brueckner-Hartree-Fock (BHF), Relativistic Brueckner (DBHF)  
variational appr., Quantum Monte Carlo  
realistic NN-interaction, no parameters
- ▶ **Effective field theory**  
Density functionals, ChPT  
perturbativ, scale arguments ( $m_\pi/M, k_F/M$ ), few parameters ( $< 2$ )
- ▶ **Empirical density functionals**  
Skyrme, Relativistic Mean Field  
many parameters (6-10), high precision fits to finite nuclei
- ▶ **Relativistic versus non-relativistic approaches**

## Overview models

- ▶ **Ab initio approaches**  
Brueckner-Hartree-Fock (BHF), Relativistic Brueckner (DBHF)  
variational appr., Quantum Monte Carlo  
realistic NN-interaction, no parameters
- ▶ **Effective field theory**  
Density functionals, ChPT  
perturbativ, scale arguments ( $m_\pi/M$ ,  $k_F/M$ ), few parameters ( $< 2$ )
- ▶ **Empirical density functionals**  
Skyrme, Relativistic Mean Field  
many parameters (6-10), high precision fits to finite nuclei
- ▶ **Relativistic versus non-relativistic approaches**



## Overview models

- ▶ **Ab initio approaches**  
Brueckner-Hartree-Fock (BHF), Relativistic Brueckner (DBHF)  
variational appr., Quantum Monte Carlo  
realistic NN-interaction, no parameters
- ▶ **Effective field theory**  
Density functionals, ChPT  
perturbativ, scale arguments ( $m_\pi/M$ ,  $k_F/M$ ), few parameters ( $< 2$ )
- ▶ **Empirical density functionals**  
Skyrme, Relativistic Mean Field  
many parameters (6-10), high precision fits to finite nuclei
- ▶ Relativistic versus non-relativistic approaches

## Overview models

- ▶ **Ab initio approaches**  
Brueckner-Hartree-Fock (BHF), Relativistic Brueckner (DBHF)  
variational appr., Quantum Monte Carlo  
realistic NN-interaction, no parameters
- ▶ **Effective field theory**  
Density functionals, ChPT  
perturbativ, scale arguments ( $m_\pi/M$ ,  $k_F/M$ ), few parameters ( $< 2$ )
- ▶ **Empirical density functionals**  
Skyrme, Relativistic Mean Field  
many parameters (6-10), high precision fits to finite nuclei
- ▶ **Relativistic versus non-relativistic approaches**

## Relativity in nuclear systems?

### Relevance of relativity:

$k_F/M \simeq 1/4 \rightarrow$  velocity  $v \simeq 1/4c$

$\rightarrow$  moderate corrections from relativistic kinematics

### But:

- ▶ Relativistic dynamics: new scale  
RMF, Hadronic many-body theory (DBHF), QCD sum rules  
 $\rightarrow \Sigma_s \simeq -350 \text{ MeV}, \Sigma_0 \simeq +300 \text{ MeV}$
- ▶ Cancellation in mean field potential  $U_{s.p.} \simeq \Sigma_o + \Sigma_s \simeq -50 \text{ MeV}$
- ▶ Large spin-orbit force  $U_{S.O.} \propto (\Sigma_o - \Sigma_s) \vec{L} \cdot \vec{S} \simeq +750 \text{ MeV}$

## Relativity in nuclear systems?

### Relevance of relativity:

$k_F/M \simeq 1/4 \rightarrow$  velocity  $v \simeq 1/4c$

$\rightarrow$  moderate corrections from relativistic kinematics

### But:

- ▶ Relativistic dynamics: new scale  
RMF, Hadronic many-body theory (DBHF), QCD sum rules  
 $\rightarrow \Sigma_s \simeq -350 \text{ MeV}, \Sigma_0 \simeq +300 \text{ MeV}$
- ▶ Cancellation in mean field potential  $U_{s.p.} \simeq \Sigma_o + \Sigma_s \simeq -50 \text{ MeV}$
- ▶ Large spin-orbit force  $U_{S.O.} \propto (\Sigma_o - \Sigma_s)\vec{L} \cdot \vec{S} \simeq +750 \text{ MeV}$

## Relativity in nuclear systems?

Known from phenomenology

boson-exchange, RMF:

large scalar/vector fields

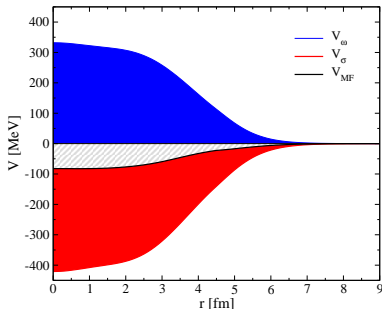
⇒ large SO force

Is the new scale universal?

relation to NN-scattering ?

large fields as a consequence of Lorentz symmetry ?

relation to chiral condensate ?



## Relativity in nuclear systems?

Known from phenomenology

boson-exchange, RMF:

large scalar/vector fields

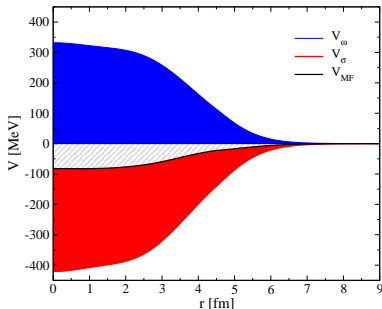
⇒ large SO force

Is the new scale universal?

relation to NN-scattering ?

large fields as a consequence of Lorentz symmetry ?

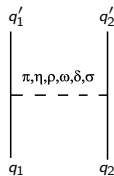
relation to chiral condensate ?



# One-boson exchange potentials

## Bonn and CD-Bonn potentials

$$V(\mathbf{q}', \mathbf{q}) = \sum_{\alpha=s, \rho S, \nu} \bar{V}_\alpha(\mathbf{q}', \mathbf{q}) \mathcal{F}_\alpha^2(\mathbf{q}', \mathbf{q}; \lambda_\alpha)$$



$$-i\bar{V}_\alpha(q', q) = \frac{\bar{u}(-\mathbf{q}') \kappa_2^{(\alpha)} u(-\mathbf{q}) P_\alpha \bar{u}(\mathbf{q}') \kappa_1^{(\alpha)} u(\mathbf{q})}{(q' - q)^2 - m_\alpha^2}, \quad u_\lambda(\mathbf{q}) = \sqrt{\frac{E+M}{2M}} \begin{pmatrix} 1 \\ \frac{2\lambda|\mathbf{q}|}{E+M} \end{pmatrix} \chi_\lambda$$

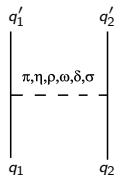
Dirac structure  $\kappa^{(s)} = g_s \mathbf{1}, \quad \kappa^{(\rho S)} = g_{\rho S} \frac{\mathbf{q}' - \mathbf{q}}{2M} i\gamma_5, \quad \kappa^{(\nu)} = g_\nu \gamma^\mu + \frac{f_\nu}{2M} i\sigma^{\mu\nu}$

→ various scales and spin-isospin structure associated with meson exchange,  
 long range=OPE, short/intermediate range = heavy mesons

# One-boson exchange potentials

## Bonn and CD-Bonn potentials

$$V(\mathbf{q}', \mathbf{q}) = \sum_{\alpha=s,ps,v} \bar{V}_\alpha(\mathbf{q}', \mathbf{q}) \mathcal{F}_\alpha^2(\mathbf{q}', \mathbf{q}; \lambda_\alpha)$$



$$-i\bar{V}_\alpha(q', q) = \frac{\bar{u}(-\mathbf{q}') \kappa_2^{(\alpha)} u(-\mathbf{q}) P_\alpha \bar{u}(\mathbf{q}') \kappa_1^{(\alpha)} u(\mathbf{q})}{(q' - q)^2 - m_\alpha^2}, \quad u_\lambda(\mathbf{q}) = \sqrt{\frac{E+M}{2M}} \begin{pmatrix} 1 \\ \frac{2\lambda|\mathbf{q}|}{E+M} \end{pmatrix} \chi_\lambda$$

Dirac structure  $\kappa^{(s)} = g_s \mathbf{1}$ ,  $\kappa^{(ps)} = g_{ps} \frac{\not{q}' - \not{q}}{2M} i\gamma_5$ ,  $\kappa^{(v)} = g_v \gamma^\mu + \frac{f_v}{2M} i\sigma^{\mu\nu}$

→ various scales and spin-isospin structure associated with meson exchange,  
 long range=OPE, short/intermediate range = heavy mesons



## Non-relativistic potentials

Low energy expansion of OBE potential

$$V(\mathbf{q}', \mathbf{q}) = \sum_{\alpha=1,5} [V_{\alpha} + V'_{\alpha} \tau_1 \cdot \tau_2] O_{\alpha}$$

$$O_1 = 1,$$

$$O_2 = \boldsymbol{\sigma}_1 \cdot \boldsymbol{\sigma}_2,$$

$$O_3 = (\boldsymbol{\sigma}_1 \cdot \mathbf{k})(\boldsymbol{\sigma}_2 \cdot \mathbf{k}),$$

$$O_4 = \frac{i}{2}(\boldsymbol{\sigma}_1 + \boldsymbol{\sigma}_2) \cdot \mathbf{n},$$

$$O_5 = (\boldsymbol{\sigma}_1 \cdot \mathbf{n})(\boldsymbol{\sigma}_2 \cdot \mathbf{n}),$$

$$\mathbf{k} = \mathbf{q}' - \mathbf{q},$$

$$\mathbf{P} = \frac{1}{2}(\mathbf{q}' + \mathbf{q}),$$

$$\mathbf{n} = \mathbf{q} \times \mathbf{q}' \equiv \mathbf{P} \times \mathbf{k},$$

## Non-relativistic potentials

Low energy expansion of OBE potential

$$V(\mathbf{q}', \mathbf{q}) = \sum_{\alpha=1,5} [V_{\alpha} + V'_{\alpha} \tau_1 \cdot \tau_2] O_{\alpha}$$

$$O_1 = 1,$$

$$O_2 = \boldsymbol{\sigma}_1 \cdot \boldsymbol{\sigma}_2,$$

$$O_3 = (\boldsymbol{\sigma}_1 \cdot \mathbf{k})(\boldsymbol{\sigma}_2 \cdot \mathbf{k}),$$

$$O_4 = \frac{i}{2}(\boldsymbol{\sigma}_1 + \boldsymbol{\sigma}_2) \cdot \mathbf{n},$$

$$O_5 = (\boldsymbol{\sigma}_1 \cdot \mathbf{n})(\boldsymbol{\sigma}_2 \cdot \mathbf{n}),$$

$$\mathbf{k} = \mathbf{q}' - \mathbf{q},$$

$$\mathbf{P} = \frac{1}{2}(\mathbf{q}' + \mathbf{q}),$$

$$\mathbf{n} = \mathbf{q} \times \mathbf{q}' \equiv \mathbf{P} \times \mathbf{k},$$

## Non-relativistic potentials

- ▶ Nijm 93 and Nijmegen I/II  
long range part due to OPE, approximate OBE amplitudes
- ▶ Argonne  $v_{18}$   
long range part due to OPE, intermediate and short range parametrized via operators  $O_\alpha$  and strength functions  $V_\alpha$
- ▶ Idaho potential  
Chiral effective field theory, N<sup>3</sup>LO, D. Entem and R. Machleidt, (29 free model parameters)
- ▶  $V_{lowk}$   
Derivation of an effective low-momentum potential  $V_{lowk}$  from modern  $NN$  potentials (out-integration of high-momentum modes,  $\Lambda \simeq 2fm^{-1}$ , and use of renormalization group methods)

## Projection onto covariant operators

$|\text{LSJ}\rangle \rightarrow$  partial wave helicity basis  $\rightarrow$  plane wave helicity basis  $\rightarrow$  Covariant basis

### Choice of basis

- ▶ **Fermi covariants**  $\Gamma_m = \{S, V, T, P, A\}$

$$S = \mathbf{1} \otimes \mathbf{1}, \quad V = \gamma^\mu \otimes \gamma_\mu, \quad T = \sigma^{\mu\nu} \otimes \sigma_{\mu\nu}, \quad P = \gamma_5 \otimes \gamma_5, \quad A = \gamma_5 \gamma^\mu \otimes \gamma_5 \gamma_\mu$$

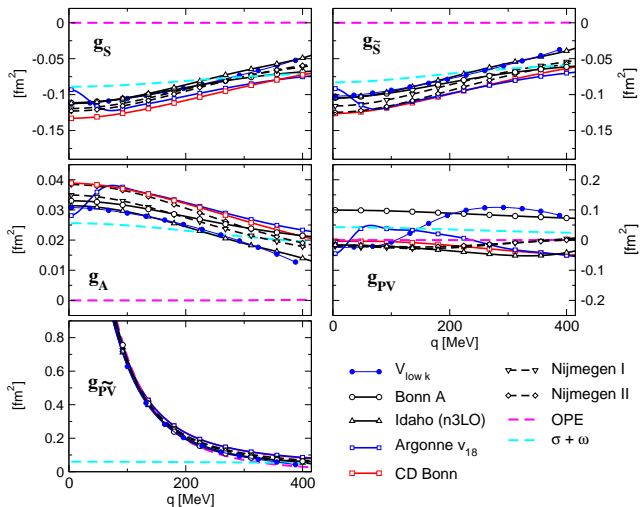
- ▶ **Pseudovector choice** (Tjon and Wallace)  $\Gamma_m = \{S, \tilde{S}, (A - \tilde{A}), PV, \tilde{P}\tilde{V}\}$

$$\text{Exchange covariants } \tilde{\Gamma}_m = \tilde{S}\Gamma_m, \quad \tilde{S}u(1)_\sigma u(2)_\tau = u(1)_\tau u(2)_\sigma$$

### Pseudovector choice

$$\begin{aligned} \hat{V}^I(|\mathbf{q}|, \theta) = & g_S^I(|\mathbf{q}|, \theta) S - g_{\tilde{S}}^I(|\mathbf{q}|, \theta) \tilde{S} + g_A^I(|\mathbf{q}|, \theta) (A - \tilde{A}) \\ & + g_{PV}^I(|\mathbf{q}|, \theta) PV - g_{\tilde{P}\tilde{V}}^I(|\mathbf{q}|, \theta) \tilde{P}\tilde{V} \end{aligned}$$

# Lorentz invariant amplitudes



## Self-energy in Hartree-Fock approximation

$$\Sigma_{\alpha\beta}(k, k_F) = -i \int_F \frac{d^4 q}{(2\pi)^4} G_{\tau\sigma}^D(q) V^A(k, q)_{\alpha\sigma;\beta\tau}$$

Dirac propagator  $G^D(q) = [\not{q} + M] 2\pi i \delta(q^2 - M^2) \Theta(q_0) \Theta(k_F - |\mathbf{q}|)$

$$\Sigma(k, k_F) = \Sigma_s(k, k_F) - \gamma_0 \Sigma_o(k, k_F) + \boldsymbol{\gamma} \cdot \mathbf{k} \Sigma_v(k, k_F),$$

$$\Sigma_s = \frac{1}{4} \int^{k_F} \frac{d^3 |\mathbf{k}|}{(2\pi)^3} \frac{M}{E_q} \left[ 4g_S - g_{\tilde{S}} + 4g_A - \frac{(k^\mu - q^\mu)^2}{4M^2} g_{\tilde{P}\tilde{V}} \right]$$

$$\Sigma_o = \frac{1}{4} \int^{k_F} \frac{d^3 |\mathbf{q}|}{(2\pi)^3} \left[ g_{\tilde{S}} - 2g_A + \frac{E_k}{E_q} \frac{(k^\mu - q^\mu)^2}{4M^2} g_{\tilde{P}\tilde{V}} \right]$$

$$\Sigma_v = \frac{1}{4} \int^{k_F} \frac{d^3 |\mathbf{q}|}{(2\pi)^3} \frac{|\mathbf{k}| \cdot |\mathbf{q}|}{|\mathbf{k}|^2 E_q} \left[ g_{\tilde{S}} - 2g_A + \frac{k_z}{q_z} \frac{(k^\mu - q^\mu)^2}{4M^2} g_{\tilde{P}\tilde{V}} \right]$$

## Self-energy in Hartree-Fock approximation

$$\Sigma_{\alpha\beta}(k, k_F) = -i \int_F \frac{d^4 q}{(2\pi)^4} G_{\tau\sigma}^D(q) V^A(k, q)_{\alpha\sigma;\beta\tau}$$

Dirac propagator  $G^D(q) = [\not{q} + M] 2\pi i \delta(q^2 - M^2) \Theta(q_0) \Theta(k_F - |\mathbf{q}|)$

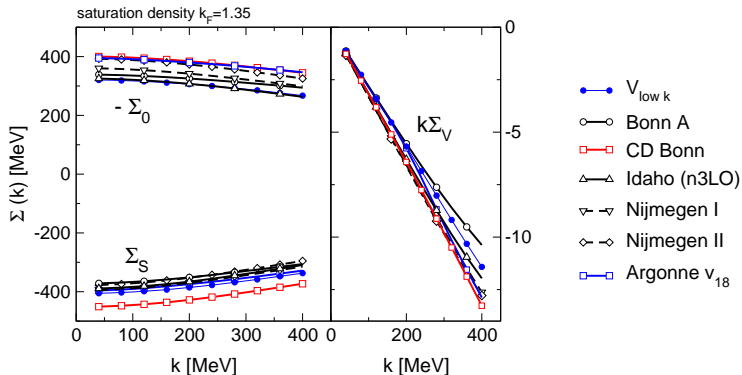
$$\Sigma(k, k_F) = \Sigma_s(k, k_F) - \gamma_0 \Sigma_o(k, k_F) + \boldsymbol{\gamma} \cdot \mathbf{k} \Sigma_v(k, k_F),$$

$$\Sigma_s = \frac{1}{4} \int^{k_F} \frac{d^3 |\mathbf{k}|}{(2\pi)^3} \frac{M}{E_{\mathbf{q}}} \left[ 4g_S - g_{\tilde{S}} + 4g_A - \frac{(k^\mu - q^\mu)^2}{4M^2} g_{\widetilde{PV}} \right]$$

$$\Sigma_o = \frac{1}{4} \int^{k_F} \frac{d^3 |\mathbf{q}|}{(2\pi)^3} \left[ g_{\tilde{S}} - 2g_A + \frac{E_{\mathbf{k}}}{E_{\mathbf{q}}} \frac{(k^\mu - q^\mu)^2}{4M^2} g_{\widetilde{PV}} \right]$$

$$\Sigma_v = \frac{1}{4} \int^{k_F} \frac{d^3 |\mathbf{q}|}{(2\pi)^3} \frac{|\mathbf{k}| \cdot |\mathbf{q}|}{|\mathbf{k}|^2 E_{\mathbf{q}}} \left[ g_{\tilde{S}} - 2g_A + \frac{k_z}{q_z} \frac{(k^\mu - q^\mu)^2}{4M^2} g_{\widetilde{PV}} \right]$$

## Large scalar/vector fields



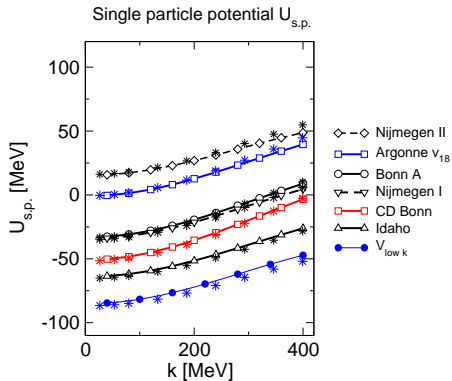
Mapping of NN potentials on relativistic operator basis

→ large scalar/vector fields → universal feature of NN interaction

O. Plohl, C.F., van Dalen, PRC 73 (2006) 014003

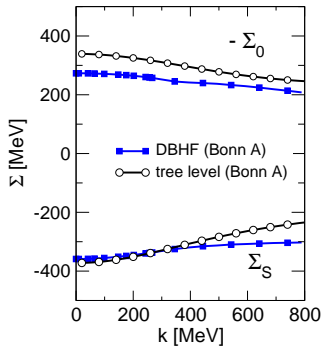


# Role of short range correlations



$$\begin{aligned}
 U_{s.p.}(k) &= \frac{M}{E_k} \langle \bar{u}(k) | \Sigma | u(k) \rangle \\
 &= M \Sigma_s / E_k - \Sigma_o + \Sigma_v k^2 / E_k
 \end{aligned}$$

Tree level HF -  
full self-consistent DBHF calculation



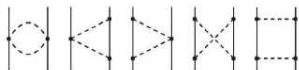
# What generates the scale in chiral EFT?

LO

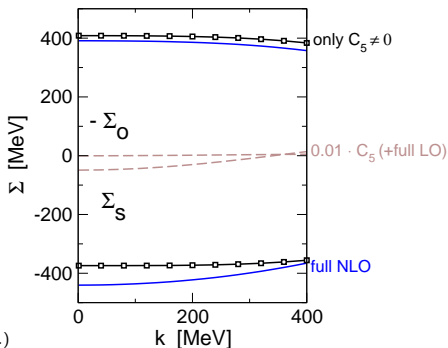


$$V = -\frac{g_A^2}{4f_\pi^2} \frac{\vec{\sigma}_1 \cdot \vec{k} \vec{\sigma}_2 \cdot \vec{k}}{q^2 + m_\pi^2}, \quad V = C_S + C_T \vec{\sigma}_1 \cdot \vec{\sigma}_2$$

NLO


 leading order  $2\pi$  exchange


$$V = \dots + C_5(-i\vec{S} \cdot (\vec{q} \times \vec{q}') + \dots + C_7(\dots))$$



Large scalar/vector fields  $\rightarrow$  NLO contact terms (strength of LEC  $C_5$  is dictated by P-wave NN scattering)

$\rightarrow$  Effective nucleon mass  $M^* = M + \Sigma_S \rightarrow$  short-distance physics

## What generates the scale in chiral EFT?

LO

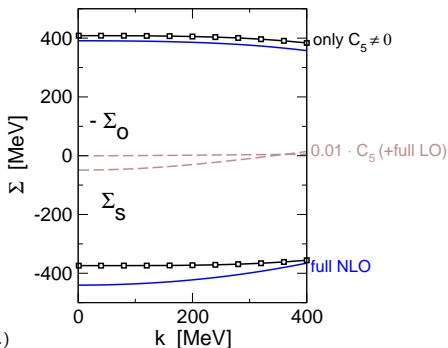


$$V = -\frac{g_A^2}{4f_\pi^2} \frac{\vec{\sigma}_1 \cdot \vec{k} \vec{\sigma}_2 \cdot \vec{k}}{q^2 + m_\pi^2}, \quad V = C_S + C_T \vec{\sigma}_1 \cdot \vec{\sigma}_2$$

NLO


 leading order  $2\pi$  exchange


$$V = \dots + C_5(-i\vec{S} \cdot (\vec{q} \times \vec{q}') + \dots + C_7(\dots))$$



Large scalar/vector fields  $\rightarrow$  NLO contact terms (strength of LEC  $C_5$  is dictated by P-wave NN scattering)

$\rightarrow$  Effective nucleon mass  $M^* = M + \Sigma_S \rightarrow$  short-distance physics

O. Ploh1, C.F., PRC 74 (2006) 034325

## Relativistic Brueckner

- ▶ N+OBEP ( $V = \sigma, \omega, \pi, \rho, \eta, \delta$ )
- ▶ 2-N correlations in hole-line expansion
- ▶ self-consistent sum of ladder diagrams

Dyson-Equation:

$$G = G_0 + G_0 \Sigma G$$



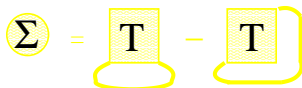
Bethe-Salpeter-Equation:

$$T = V + i \int V G G Q T$$



Self Energy:

$$\Sigma(\rho, k) = \sum_{q \in F} \langle q | T(q, k) | q \rangle = \Sigma_S - \gamma_0 \Sigma_0 + \vec{\gamma} \cdot \vec{k} \Sigma_V$$



## Technicalities/approximations

- ▶ 3-dim. reduction of BSE: Thompson eq.
- ▶ angle averaged Pauli operator
- ▶ Quadratic approx. to s.p. potential:  $\Sigma_{S,0}(k_F)$

Brockmann/Machleidt 90, Envik et al., Sammarunca et al.,...

$$U_{s.p.}(k) = \frac{m^*}{E^*} \sum_{q \in F} \langle kq | T(q, k) | kq \rangle = \frac{m^* \Sigma_S}{\sqrt{k^2 + (M + \Sigma_S)^2}} - \Sigma_0$$

- ▶ Projection on covariant amplitudes:  $\Sigma_{S,0,V}(k_F, k)$

Horowitz/Serot 87, Malfliet et. al., Tübingen

- ▶ Include negative energy states

Weigel et al., DeJong/Lenske

## PV choice reduces on-shell ambiguities

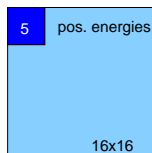
- ▶ On-shell: Equal  $ps$  and  $p\nu$  matrix elements
- ▶  $ps$  couples to negative energy states:

$$\langle \bar{u}|P|u \rangle = \langle \bar{u}|PV|u \rangle, \quad \langle \bar{v}|P|u \rangle = 1, \quad \langle \bar{v}|PV|u \rangle = 0$$

- ▶ Different contributions to  $\Sigma$ :

$$\text{tr}[\tilde{P}G(q)] \gg \text{tr}[\tilde{P}V G(q)]$$

- ▶ critical for  $1-\pi$ -exchange:  
 $p\nu$  coupling removes spurious contributions to  $\Sigma_S$  and  $\Sigma_0$
- ▶ Only possible in Tjon&Wallace basis



## Isospin asymmetric matter

- ▶ Coupled channel problem for the  $np$  channel
- ▶ 6 instead of 5 independent helicity/covariant amplitudes

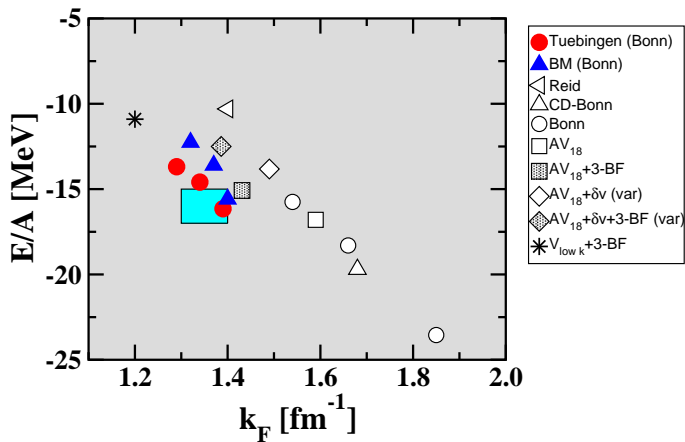
$$T_{nn} = V_{nn} + i \int V_{nn} Q_{nn} G_n G_n T_{nn}$$

$$T_{pp} = V_{pp} + i \int V_{pp} Q_{pp} G_p G_p T_{pp}$$

$$T_{np}^D = V_{np}^D + i \int V_{np}^D Q_{np} G_n G_p T_{np}^D + i \int V_{np}^X Q_{pn} G_p G_n T_{np}^X$$

$$T_{np}^X = V_{np}^X + i \int V_{np}^X Q_{pn} G_p G_n T_{np}^D + i \int V_{np}^D Q_{np} G_n G_p T_{np}^X$$

## Model comparison: nuclear matter



see e.g. C.F. arXiv:0711.3367



## Bulk properties (with Bonn A)

- ▶ Saturation properties:

$$\rho_{sat} = 0.184 \text{ fm}^{-3} \quad E_B = -16.15 \quad K = 230 \text{ MeV} \quad M^* = 637 \text{ MeV}$$

- ▶ Symmetry energy:

$$E_{sym} = 31.6 \text{ MeV} \quad @ \quad \rho = 0.160 \text{ fm}^{-3}$$

- ▶ Maximal neutron star mass:

$$M = 2.33 M_{\odot}$$

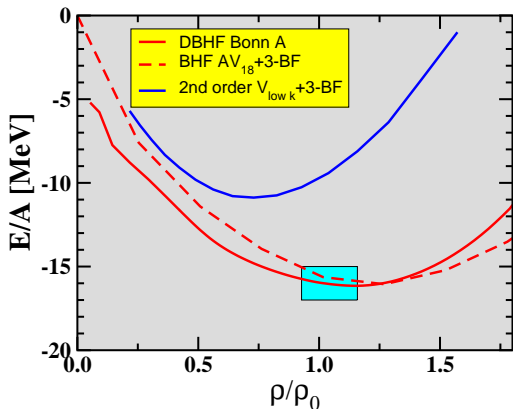
Gross-Boelting, C.F., Faessler, NPA 648 (1999) 105

van Dalen, C.F., Faessler, NPA 744 (2004) 227

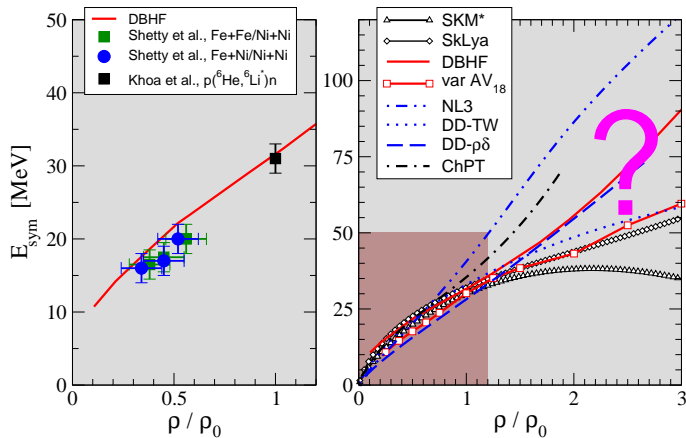
## Model comparison: EOS

Perturbative treatment with  $V_{low\ k}+3\text{-BFs}$  (Bogner et al, NPA 763 (2005) 59)

BHF with  $AV_{18}+3\text{-BFs}$  (Catania group)



## Model comparison: symmetry energy



$$E(\rho, \beta) = E(\rho) + E_{\text{sym}}(\rho)\beta^2 + \mathcal{O}(\beta^4) + \dots \quad \beta = Y_n - Y_p$$

## Effective nucleon mass

Different definitions are used!

- ▶ Non-relativistic mass:

$$m_{NR}^* = \left[ M + \frac{1}{k} \frac{d}{dk} U_{s.p.} \right]^{-1} = |\mathbf{k}| [dE/d|\mathbf{k}|]^{-1}$$

parameterizes non-locality in space (k-mass) and time (e-mass)

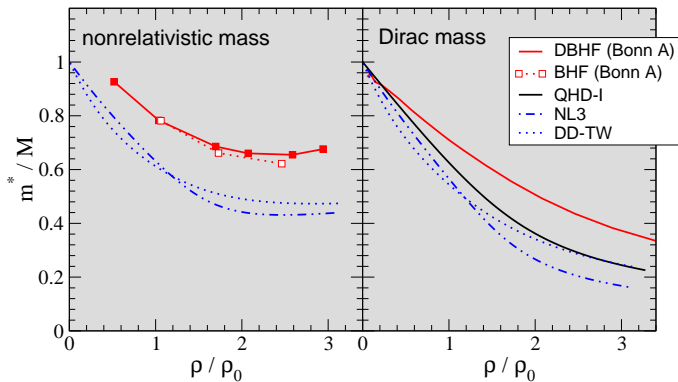
Mahaux et al., M uther, Frick,..

- ▶ Dirac mass:

$$m_D^* = M + \Sigma_S \quad , \quad U_{s.p.} \simeq \frac{m_D^*}{E^*} \Sigma_S + \Sigma_0$$

scalar part of self energy

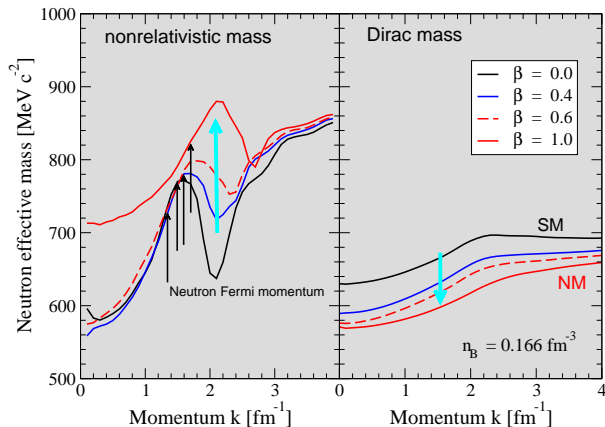
## Effective nucleon mass



## Proton-neutron mass splitting

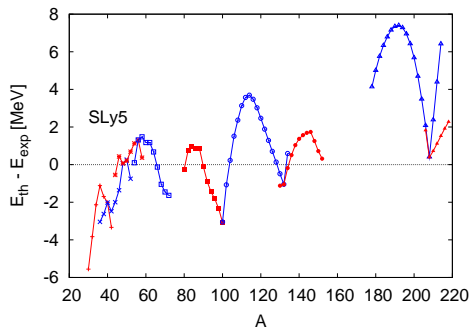
- ▶ BHF:  $m_{NR,n}^* > m_{NR,p}^*$
- ▶ RMF:  $m_{D,n}^* < m_{D,p}^*$  ;  $m_{NR,n}^* < m_{NR,p}^*$  ( $\rho + \delta$ )  
Baran, Di Toro et al., Phys. Rep. 410 ('05) 335
- ▶ DBHF with  $\Sigma$  extracted by fit method:  $m_{D,n}^* > m_{D,p}^*$   
Alonso & Sammarunca, PRC 67 ('03) 054301
- ▶ DBHF with projection method:  $m_{D,n}^* < m_{D,p}^*$   
de Jong & Lenske, PRC 58 ('98) 890; van Dalen, C.F., Faessler, NPA 744 ('04) 227
- ▶ non-rel. mass in DBHF:  $m_{NR,n}^* > m_{NR,p}^*$   
van Dalen, C.F, Faessler, PRL 95 (2005) 022302

# Proton-neutron mass splitting



## Guidance for phenomenology

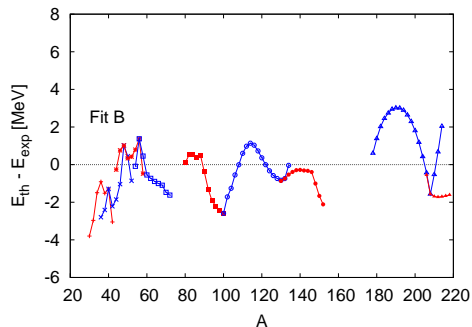
SkLy5: mass splitting  $m_n^* < m_p^*$  (in contrast to BHF/DBHF)





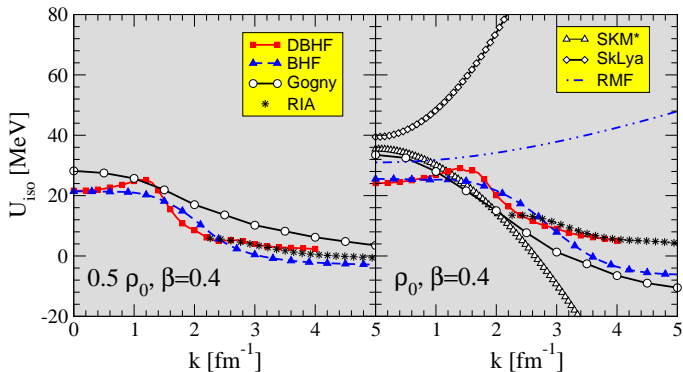
## Guidance for phenomenology

SkLy5: mass splitting  $m_n^* < m_p^*$  (in contrast to BHF/DBHF)



new fit, small change of parameters:  $m_n^* > m_p^*$  ( $\kappa_V = 0.3$ )

Lesinski et al., PRC 74 (2006) 044315

Isovector optical potential:  $U_{iso} = (U_n - U_p)/2\beta$ 

DBHF: van Dalen, C.F., Faessler, PRC 72 (2005) 065803

BHF: Zuo et al. PRC 72 (2005) 014005

RIA: Chen, Ko, Li, PRC 72 (2005) 064606

## Main differences between DBHF/BHF

- ▶ vector/scalar density

$$\rho_B \propto a^\dagger a - b^\dagger b$$

$$\rho_S \propto a^\dagger a + b^\dagger b$$

- ▶ dressed interaction:

$$\langle u|V|u \rangle \mapsto \langle u^*|V|u^* \rangle$$

- ▶ quenching of tensor force by  $m^*/M$   
iterated OPE is less important for saturation
- ▶ New scale: large scalar/vector fields
- ▶ effective inclusion of 3-body forces:  
box diagrams, intermediate  $\Delta$   
Z-graphs

## Main differences between DBHF/BHF

- ▶ vector/scalar density

$$\rho_B \propto a^\dagger a - b^\dagger b$$

$$\rho_S \propto a^\dagger a + b^\dagger b$$

- ▶ dressed interaction:

$$\langle u | V | u \rangle \mapsto \langle u^* | V | u^* \rangle$$

- ▶ quenching of tensor force by  $m^*/M$   
iterated OPE is less important for saturation
- ▶ New scale: large scalar/vector fields
- ▶ effective inclusion of 3-body forces:  
box diagrams, intermediate  $\Delta$   
Z-graphs

## Main differences between DBHF/BHF

- ▶ vector/scalar density

$$\rho_B \propto a^\dagger a - b^\dagger b$$

$$\rho_S \propto a^\dagger a + b^\dagger b$$

- ▶ dressed interaction:

$$\langle u | V | u \rangle \mapsto \langle u^* | V | u^* \rangle$$

- ▶ quenching of tensor force by  $m^*/M$   
iterated OPE is less important for saturation
- ▶ New scale: large scalar/vector fields
- ▶ effective inclusion of 3-body forces:  
box diagrams, intermediate  $\Delta$   
Z-graphs

## Main differences between DBHF/BHF

- ▶ vector/scalar density

$$\rho_B \propto a^\dagger a - b^\dagger b$$

$$\rho_S \propto a^\dagger a + b^\dagger b$$

- ▶ dressed interaction:

$$\langle u|V|u \rangle \mapsto \langle u^*|V|u^* \rangle$$

- ▶ quenching of tensor force by  $m^*/M$   
iterated OPE is less important for saturation
- ▶ New scale: large scalar/vector fields
- ▶ effective inclusion of 3-body forces:  
box diagrams, intermediate  $\Delta$   
Z-graphs

## Main differences between DBHF/BHF

- ▶ vector/scalar density

$$\rho_B \propto a^\dagger a - b^\dagger b$$

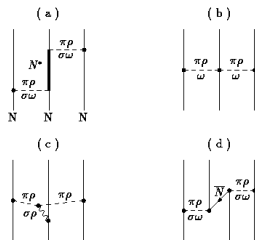
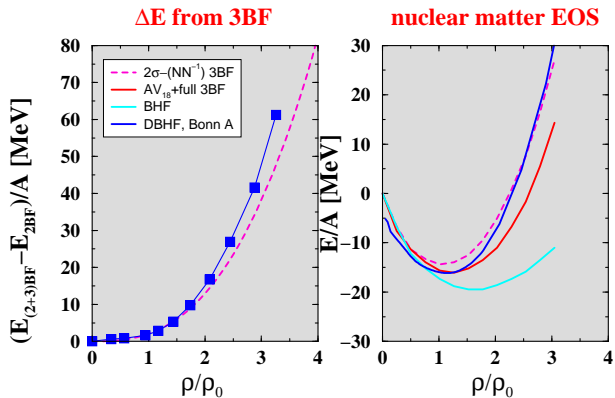
$$\rho_S \propto a^\dagger a + b^\dagger b$$

- ▶ dressed interaction:

$$\langle u|V|u \rangle \mapsto \langle u^*|V|u^* \rangle$$

- ▶ quenching of tensor force by  $m^*/M$   
iterated OPE is less important for saturation
- ▶ New scale: large scalar/vector fields
- ▶ effective inclusion of 3-body forces:  
box diagrams, intermediate  $\Delta$   
Z-graphs

# 3-body forces

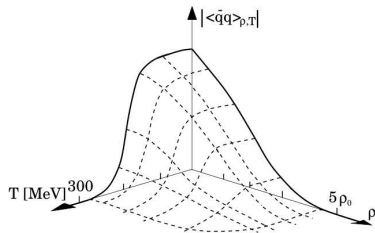
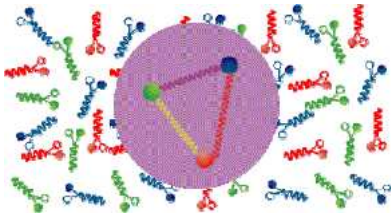




## Connection to QCD

- ▶ scalar condensate:  $\langle \bar{q}q \rangle \sim (-250 \text{ MeV})^3$
- ▶ naive QCD sum rule interpretation:

$$\frac{\langle \bar{q}q \rangle_{\rho_B}}{\langle \bar{q}q \rangle_0} = \frac{M^*}{M}$$



## In-medium chiral condensate

Hellmann-Feynman:

$$\frac{\langle \bar{q}q \rangle_{\rho_N}}{\langle \bar{q}q \rangle_0} = 1 - \frac{\rho_N}{m_\pi^2 f_\pi^2} \left[ \sigma_N + m \frac{d}{dm} \frac{E}{A} \right]$$

pion-nucleon sigma-term  $\sigma_N = m \frac{dM}{dm} = \langle N | m \bar{q}q | N \rangle$

Problem: unknown quark mass dependence

$$\sum_{\sigma, \omega, \pi, \rho} \left[ \frac{\partial E}{\partial m_i} \frac{dm_i}{dm} + \frac{\partial E}{\partial g_i} \frac{dg_i}{dm_i} + \dots \right] = ???$$

see e.g. Brockmann, Weise, PLB 367 (1996)

## In-medium chiral condensate

Hellmann-Feynman:

$$\frac{\langle \bar{q}q \rangle_{\rho N}}{\langle \bar{q}q \rangle_0} = 1 - \frac{\rho N}{m_\pi^2 f_\pi^2} \left[ \sigma_N + m \frac{d}{dm} \frac{E}{A} \right]$$

pion-nucleon sigma-term  $\sigma_N = m \frac{dM}{dm} = \langle N | m \bar{q}q | N \rangle$

Problem: unknown quark mass dependence

$$\sum_{\sigma, \omega, \pi, \rho} \left[ \frac{\partial E}{\partial m_i} \frac{dm_i}{dm} + \frac{\partial E}{\partial g_i} \frac{dg_i}{dm_i} + \dots \right] = ???$$

see e.g. Brockmann, Weise, PLB 367 (1996)

## In-medium chiral condensate

### Solution:

Determine  $E$  from chiral EFT NN interaction

Quark mass dependence known up to NLO

Epelbaum, Glöckle, Meissner, EPJA **18**, 499 (2003)

### Exact calculation at NLO

Hartree-Fock

Brückner-Hartree-Fock

O. Ploh1, C.F., NPA 798 (2008) 75

## In-medium chiral condensate

### Solution:

Determine  $E$  from chiral EFT NN interaction

Quark mass dependence known up to NLO

Epelbaum, Glöckle, Meissner, EPJA **18**, 499 (2003)

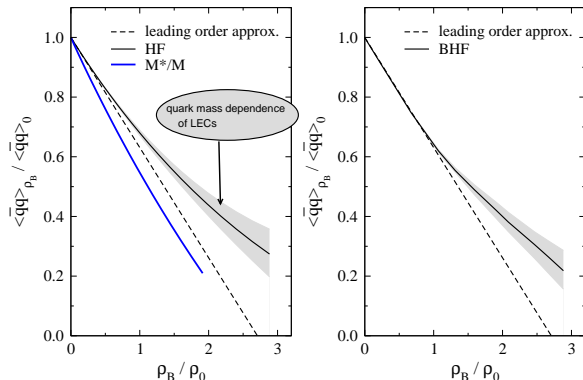
### Exact calculation at NLO

Hartree-Fock

Brückner-Hartree-Fock

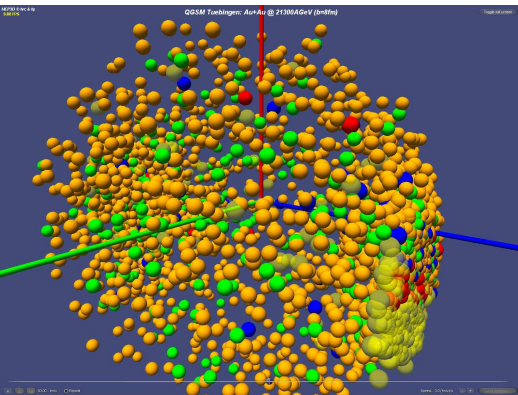
O. Ploh1, C.F., NPA 798 (2008) 75

## Chiral condensate at NLO



- ▶ decoupling of condensate and effective mass
- ▶ condensate driven by long-distance physics
- ▶ effective mass driven by short-distance physics

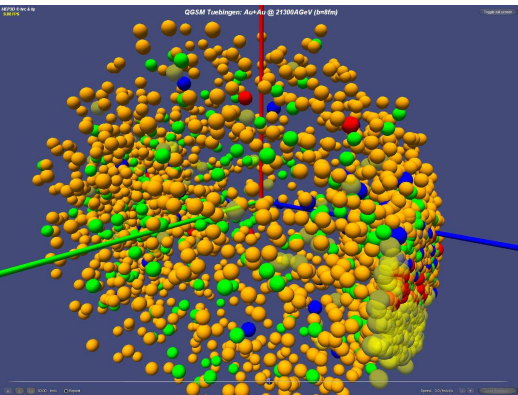
## Probing dense matter by HICs



$Au + Au$ ,  $\sqrt{s} = 200$  AGeV, QGSM Tübingen

- ▶ **Advantage:**  
On earth experiments
- ▶ **Disdvantages:**  
Short time scale, non-equilibrium  
Heated matter  
Small asymmetries  
Surface effects
- ▶ **Observables:**  
Collective flow  
Particle production

## Probing dense matter by HICs

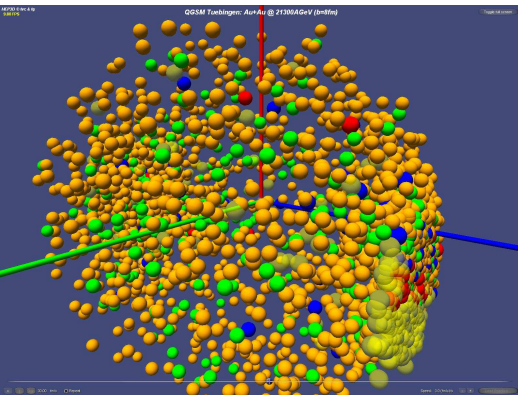


$Au + Au, \sqrt{s} = 200 \text{ AGeV}, \text{QGSM Tübingen}$

- ▶ **Advantage:**  
On earth experiments
- ▶ **Disdvantages:**  
Short time scale, non-equilibrium  
Heated matter  
Small asymmetries  
Surface effects
- ▶ **Observables:**  
Collective flow  
Particle production



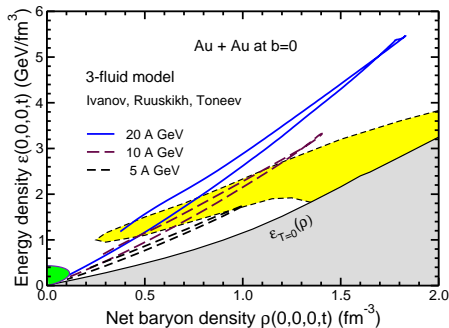
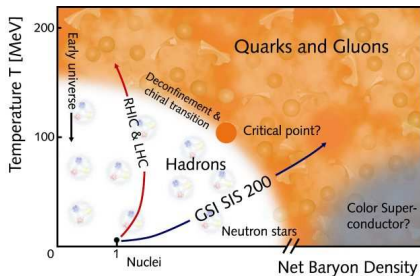
## Probing dense matter by HICs



$Au + Au$ ,  $\sqrt{s} = 200$  AGeV, QGSM Tübingen

- ▶ **Advantage:**  
On earth experiments
- ▶ **Disadvantages:**  
Short time scale, non-equilibrium  
Heated matter  
Small asymmetries  
Surface effects
- ▶ **Observables:**  
Collective flow  
Particle production

## Densities at CBM/FAIR



- ▶ DOFs, phase transition?
- ▶ SIS: symmetric nuclear matter up to  $3 \rho_0$

see e.g. CBM Physics Book: Collision Dynamics

## Subthreshold kaon production

- ▶ Subthreshold particle production:  $E_{Lab} < E_{thr}$ :
 

$K^+(u\bar{s})$ :	$NN \mapsto N\Lambda K^+$	$E_{thr} = 1.58 \text{ GeV}$
	$\pi N \mapsto \Lambda K^+$	
$K^-(\bar{u}s)$ :	$NN \mapsto NNK^+K^-$	$E_{thr} = 2.5 \text{ GeV}$
- ▶ Energy provided by multistep collisions:  
excludes large surface effects
- ▶  $K^+(u\bar{s})$ : strangeness conservation:  
(almost) no final state interaction
- ▶ ideal probe for dense phase

## Subthreshold kaon production

- ▶ Subthreshold particle production:  $E_{Lab} < E_{thr}$ :
 

$K^+(u\bar{s})$ :	$NN \mapsto N\Lambda K^+$	$E_{thr} = 1.58 \text{ GeV}$
	$\pi N \mapsto \Lambda K^+$	
$K^-(\bar{u}s)$ :	$NN \mapsto NNK^+K^-$	$E_{thr} = 2.5 \text{ GeV}$
- ▶ Energy provided by multistep collisions:  
excludes large surface effects
- ▶  $K^+(u\bar{s})$ : strangeness conservation:  
(almost) no final state interaction
- ▶ ideal probe for dense phase

## Subthreshold kaon production

- ▶ Subthreshold particle production:  $E_{Lab} < E_{thr}$ :
 

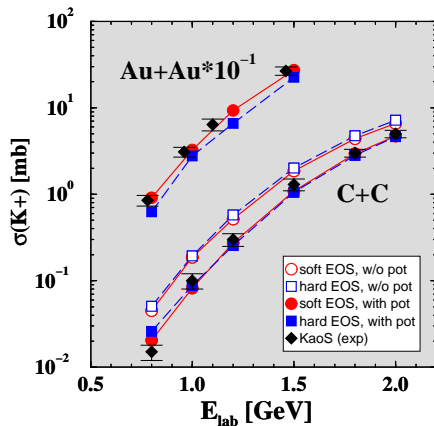
$K^+(u\bar{s})$ :	$NN \mapsto N\Lambda K^+$	$E_{thr} = 1.58 \text{ GeV}$
	$\pi N \mapsto \Lambda K^+$	
$K^-(\bar{u}s)$ :	$NN \mapsto NNK^+K^-$	$E_{thr} = 2.5 \text{ GeV}$
- ▶ Energy provided by multistep collisions:  
excludes large surface effects
- ▶  $K^+(u\bar{s})$ : strangeness conservation:  
(almost) no final state interaction
- ▶ ideal probe for dense phase

## Subthreshold kaon production

- ▶ Subthreshold particle production:  $E_{Lab} < E_{thr}$ :
 

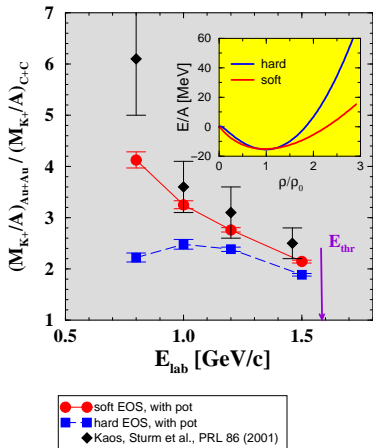
$K^+(u\bar{s})$ :	$NN \mapsto N\Lambda K^+$	$E_{thr} = 1.58 \text{ GeV}$
	$\pi N \mapsto \Lambda K^+$	
$K^-(\bar{u}s)$ :	$NN \mapsto NNK^+K^-$	$E_{thr} = 2.5 \text{ GeV}$
- ▶ Energy provided by multistep collisions:  
excludes large surface effects
- ▶  $K^+(u\bar{s})$ : strangeness conservation:  
(almost) no final state interaction
- ▶ ideal probe for dense phase

## RQMD transport calculations



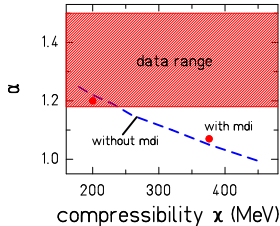
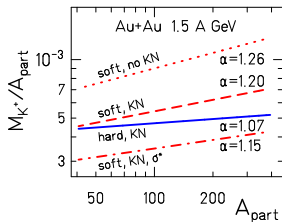
Data: Sturm et al., KaoS Coll., PRL 86 (2001) 39

## Conclusions from KaoS data



C.F. et al., PRL 86 (2001) 1974

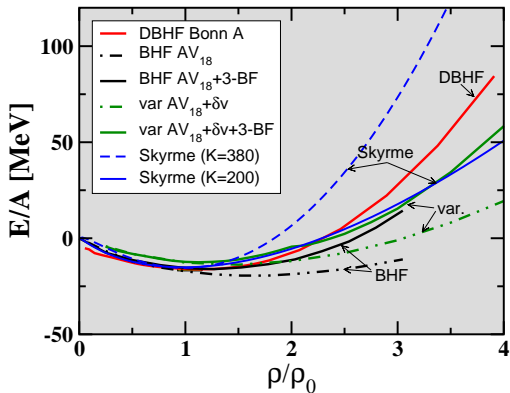
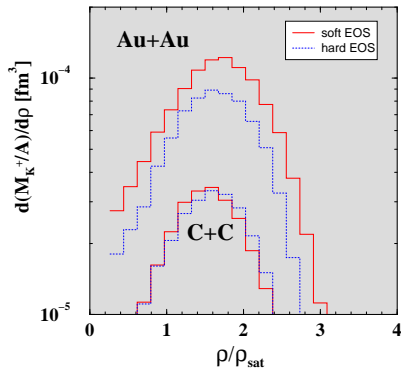
C.F., Prog. Part. Nucl. Phys. 56 (2006) 1



Hartnack et al. PRL 96 (2006) 012302



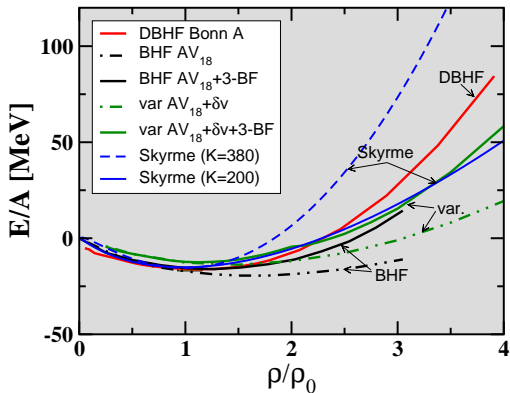
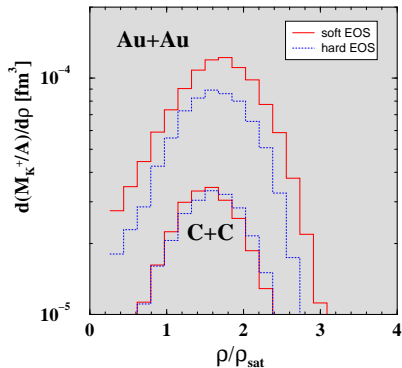
## Density range tested by $K^+$ @ SIS



e.g. Au+Au @ 0.8 AGeV  $\langle \rho/\rho_0 \rangle = 1.53$

Below  $2.5 \div 3\rho_0$  EOS is soft!

## Density range tested by $K^+$ @ SIS

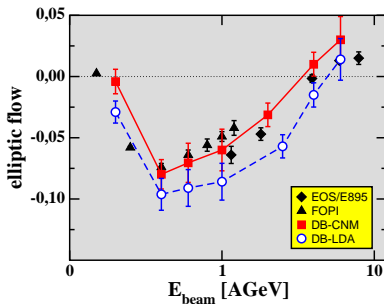
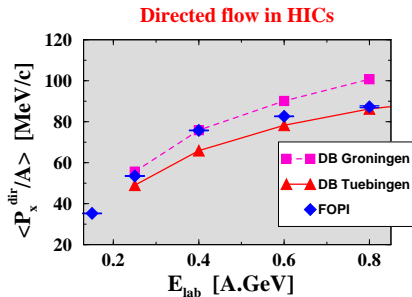


e.g. Au+Au @ 0.8 AGeV  $\langle \rho/\rho_0 \rangle = 1.53$

Below  $2.5 \div 3\rho_0$  EOS is soft!

## Directed/elliptic flow

RBUU transport: e.g. Gaitanos et al., EPJA 12 (2001) 421; C.F., Gaitanos, NPA 714 (2003) 634



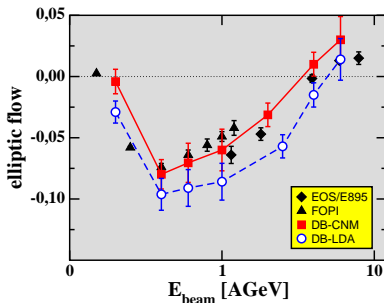
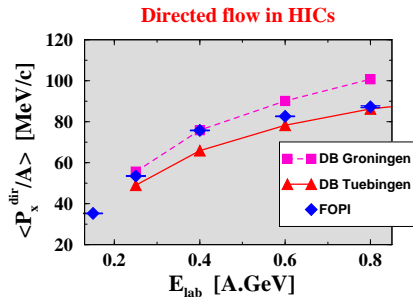
Large non-equilibrium effects:

→ makes the link to equilibrium EOS difficult → careful treatment!

→ fair agreement with data

## Directed/elliptic flow

RBUU transport: e.g. Gaitanos et al., EPJA 12 (2001) 421; C.F., Gaitanos, NPA 714 (2003) 634



Large non-equilibrium effects:

→ makes the link to equilibrium EOS difficult → careful treatment!

→ fair agreement with data

## Summary & conclusions

- ▶ Relativity
  - new scale: large fields
  - direct consequence of NN-force (SO-force)
- ▶ many-body-theory: DBHF
  - “best- bulk properties
  - correct pn mass-splitting
  - isovector potential
- ▶ EOS
  - soft at moderate, stiff at high densities
  - asi-stiff
- ▶ decoupling of effective mass and scalar condensate
  - short-distance / long-distance physics
- ▶ DBHF EOS in agreement with
  - heavy ion collisions
  - astrophysical constraints

## Summary & conclusions

- ▶ Relativity
  - new scale: large fields
  - direct consequence of NN-force (SO-force)
- ▶ many-body-theory: DBHF
  - “best- bulk properties
  - correct pn mass-splitting
  - isovector potential
- ▶ EOS
  - soft at moderate, stiff at high densities
  - asi-stiff
- ▶ decoupling of effective mass and scalar condensate
  - short-distance / long-distance physics
- ▶ DBHF EOS in agreement with
  - heavy ion collisions
  - astrophysical constraints

## Summary & conclusions

- ▶ Relativity
  - new scale: large fields
  - direct consequence of NN-force (SO-force)
- ▶ many-body-theory: DBHF
  - “best- bulk properties
  - correct pn mass-splitting
  - isovector potential
- ▶ EOS
  - soft at moderate, stiff at high densities
  - asi-stiff
- ▶ decoupling of effective mass and scalar condensate
  - short-distance / long-distance physics
- ▶ DBHF EOS in agreement with
  - heavy ion collisions
  - astrophysical constraints

## Summary & conclusions

- ▶ Relativity
  - new scale: large fields
  - direct consequence of NN-force (SO-force)
- ▶ many-body-theory: DBHF
  - “best- bulk properties
  - correct pn mass-splitting
  - isovector potential
- ▶ EOS
  - soft at moderate, stiff at high densities
  - asi-stiff
- ▶ decoupling of effective mass and scalar condensate
  - short-distance / long-distance physics
- ▶ DBHF EOS in agreement with
  - heavy ion collisions
  - astrophysical constraints



## Summary & conclusions

- ▶ Relativity
  - new scale: large fields
  - direct consequence of NN-force (SO-force)
- ▶ many-body-theory: DBHF
  - “best- bulk properties
  - correct pn mass-splitting
  - isovector potential
- ▶ EOS
  - soft at moderate, stiff at high densities
  - asi-stiff
- ▶ decoupling of effective mass and scalar condensate
  - short-distance / long-distance physics
- ▶ DBHF EOS in agreement with
  - heavy ion collisions
  - astrophysical constraints