⁵He ternary fission yields of 252 Cf and 235 U(n, f)

J. K. Hwang,¹ A. V. Ramayya,¹ J. H. Hamilton,¹ C. J. Beyer,¹ J. Kormicki,¹ X. Q. Zhang,¹ A. Rodin,² A. Formichev,² J. Kliman,² L. Krupa,² G. M. Ter Akopian,² Yu. Ts. Oganessian,² G. Hubarian,³ D. Seweryniak,⁴ C. J. Lister,⁴ R. V. F. Janssens,⁴ I. Ahmad,⁴ M. P. Carpenter,⁴ J. P. Greene,⁴ T. Lauritsen,⁴ I. Wiedenhöver,⁴ W. C. Ma,⁵ R. B. Piercey,⁵

and J. D. Cole⁶

¹Department of Physics, Vanderbilt University, Nashville, Tennessee 37235

²Flerov Laboratory for Nuclear Reactions, Joint Institute for Nuclear Research, Dubna, Russia

³Cyclotron Institute, Texas A & M University, Texas 77843-3366

⁴Argonne National Laboratory, Argonne, Illinois 60439

⁵Department of Physics, Mississippi State University, Mississippi 39762

⁶Idaho National Engineering and Environmental Laboratory, Idaho Falls, Idaho 83415

(Received 8 September 1999; published 15 March 2000)

The relative ⁴He and ⁵He ternary fission yields were determined from a careful analysis of the energy distribution of α spectra from a new measurement with a ²⁵²Cf source and from published data on ²⁵²Cf and 235 U(*n*,*f*). The kinetic energies of the ⁵He and ⁴He ternary particles were found to be approximately 11 and 16 MeV, respectively. ⁵He particles contribute 10–20 % to the total alpha yield with the remainder originating from ⁴He accompanied fission.

PACS number(s): 25.85.Ca, 27.90.+b, 29.30.Ep

Since the discovery of particle-accompanied fission by Alvarez in 1943 [1], the study of ternary spontaneous fission has been of much interest. This process is a rare one as it occurs about once in \sim 500 spontaneous fission events. A helium atom is present in over 90% of all ternary fission events. In the case of ²⁵²Cf, the relative yields of ⁴He, ⁶He, and ⁸He have been measured to be 10^4 , 393(60), and 25(5), respectively, with an absolute ternary ⁴He yield of 3.82×10^{-3} [2,3]. The contribution of ⁵He to the ternary helium fission yield is difficult to determine because of its breakup into the $\alpha + n$ channel with a half-life of 8 $\times 10^{-22}$ sec. The Q value and neutron energy associated with this process are 0.957 MeV and 4.0(3) MeV, respectively [4], e.g., $\langle E_{^{5}\text{He}} \rangle = \langle E_{\alpha} \rangle + 3$ (MeV). The presence of ⁵He in the fission of ²⁵²Cf was originally demonstrated by Cheifetz et al. [4], who measured the correlation between neutrons and α particles at 0° and 180°, and reported that approximately 11(2)% of the ternary α particles result from ⁵He breakup. The average ⁵He kinetic energy was measured to be 15.4 MeV, i.e., the associated α energy is 12.4(9) MeV [4]. The latter value should be contrasted with the *average* α energy of 15.7 MeV reported in Ref. [2].

In principle, an indirect estimate of the ternary ⁵He fission yield can be derived from a calculation of the mean extra energy cost (E_c) required to emit a ternary particle [5]. This results in a $\sim 5\%$ contribution to the total ternary α fission yield for $^{235}U(n,f)$. However, small changes in E_c impact the predicted yields significantly, and this approach can only serve as a rough guide [5,6]. A third, and possibly more reliable, estimate of the ternary ⁵He yield comes from an attempt to understand the shape of the energy spectrum of the ternary α particles [6]. Remarkably, the energy spectra of all ternary particles are Gaussian (this observation is independent of the excitation energy of the fissioning system), except that of the α particles. In this case there is a lowenergy tail which can be attributed to α particles originating from ⁵He emission. The contributions from the breakup of the ⁶He and ⁸He can be neglected in this analysis since these processes will be less than 1%. In Refs. [6,7] this low-energy tail was established to correspond to a mean energy of ~ 6 MeV and the associated yield was estimated to be 15.5% of the total α ternary particles for ²⁵²Cf [7].

The results of Refs. [4-7] discussed above leave a number of open questions. The mean ⁵He energy of ~ 6 MeV extracted from the low-energy tail of the α spectra [6,7] differs significantly from the only direct measurement of 12.4 MeV [4]. Furthermore, the ⁵He yield extracted from this tail, 15.5% [6,7], is barely consistent with the yield of 11% reported by Cheifetz et al. [4]. Additional uncertainties regarding the experimental situation come from the fact that the role of the neutron multiplicity was not considered in Ref. [4]. This multiplicity was determined later by Han Hongyin et al. [8]. Thus, it appeared worthwhile to revisit the issue. Hereafter, we present the results of a new analysis of α spectra from ternary fission. Data from a new measurement with a ²⁵²Cf source as well as published data on ²⁵²Cf and $^{235}U(n,f)$ have been analyzed. A consistent picture of the ⁵He ternary fission yield appears to emerge.

The energy spectrum of α particles emitted in the spontaneous fission of 252 Cf was measured by using two ΔE -E Si detector telescopes installed at the center of the Gammasphere array [9] at Argonne National Laboratory. The ²⁵²Cf source with a strength of the order of 30 μ Ci was sandwiched between two Au foils, each 19 mg/cm² in thickness. These foils were used to stop the fission fragments so that problems associated with the Doppler shift of prompt fission γ rays were minimized. In addition, on each side of the source a 0.84 mg/cm² Mylar foil was added to provide full absorption of α particles emitted in the radioactive decay of ²⁵²Cf. This approach limited potential radiation damage to the Si detectors. Each ΔE -E telescope included one ΔE detector (of 15 μ m thickness and 10×10 mm² area) and one BRIEF REPORTS



strip detector. The strip detectors were 400 μ m and 60 $\times 60 \text{ mm}^2$ in thickness and area, respectively and had 29 strips. The 2 mm strips were connected in pairs so that, for each strip detector, 15 signals could be obtained from 14 pairs and one single strip. This implied a 4 mm position resolution along the x axes of each strip detector, whereas there was no position resolution along the y axes. With the ΔE and E (strip) telescopes installed at distances of 6 and 40 mm from the source, respectively, the condition was fulfilled that all light charged particles detected by either of the strip detectors passed through the corresponding ΔE detector and gave ΔE signals. With the position resolution of the strip detector (4 mm wide strips and 1 mm resolution along each strip), the ΔE -E telescopes provided unambiguous Z and A identification for all the light charged particles of interest. The energy calibration of the telescopes was performed with ²²⁴Ra and ²²⁸Th radioactive sources and with the 6.2 MeV alpha peak from the decay of ²⁵²Cf deposited on the surface of the Au foil (see below). This calibration automatically takes the presence of the Mylar foil into account. A software correction was applied for the energy loss in the Au foil. The latter was checked as an internal energy calibration was also obtained by using the average energy of 15.7 MeV of the broad ternary α peak as reported by Mutterer *et al.* [2].



The typical α spectrum from the present measurement is given in Fig. 1. As expected, a broad energy distribution was recorded. The small peak visible at 6.2 MeV comes from the α decay of ²⁵²Cf. Its presence in the data is due to the accidental deposition of a small amount of source material on the surface of the Au foil. The observed mean energy of 13.1 MeV agrees well with the expected energy of 15.7 MeV [2], once the energy loss through the Au fold is taken into account. The α spectrum for the subsequent analysis, i.e., the spectrum corrected for energy loss and with the 6.2 MeV contamination removed, is shown in Fig. 2. No reliable data could be obtained below 9.0 MeV, the low energy cutoff of the measurement.

In order to disentangle the respective contributions of ⁵He and ⁴He to the alpha particle spectrum observed with the ²⁵²Cf source, attempts were made to fit the data with two Gaussian functions. Two methods were used. In the first of these, the energy of the ⁵He particles was kept fixed at the energy reported by Cheifetz *et al.* in Ref. [4], while all other parameters defining the two Gaussians were varied. In particular, in this approach the ⁴He energy is a free parameter. The results are tabulated in the top rows of Table I and the quality of the fit can be judged from Fig. 2. The extracted yield for ternary ⁵He emission is ~35% of the total α yield,



FIG. 2. The α energy distribution with the correction for the 6.2 MeV α decay and the energy loss through the Au foil.

TABLE I. Yield ratios extracted for the ternary ⁴He and ⁵He particles of ²⁵²Cf in the present work. $\langle E_{^5\text{He}} \rangle = \langle E_{\alpha} \rangle - Q + \langle E_n \rangle = \langle E_{\alpha} \rangle + 3$ (MeV).

	Ternary particle	$\langle E \rangle$ MeV	FWHM MeV	Yield ratio	χ^2 /DOF
Method 1	⁴ He	$16.3^{+0.3}_{-0.3}$	$7.3^{-0.3}_{+0.1}$	$65.2^{-2.8}_{+2.8}$	0.065
	⁵ He	15.4[3]	$8.9^{-1.2}_{+1.3}$	$34.8^{+1.4}_{-1.8}$	
Method 2	⁴ He	$16.0^{+1.0}_{-0.5}$	$7.4^{-0.8}_{+0.4}$	$79.6^{-24.4}_{+12.0}$	0.068
	⁵ He	$13.9^{+1.1}_{-4.3}$	$7.4^{-0.8}_{+0.4}$	$20.4^{+21.9}_{-2.9}$	

i.e., this process occurs roughly half as often as ternary fission involving ⁴He. Hence, the ⁵He intensity derived in this way is larger than that reported in Ref. [4] (11%). However, in the fitting process, the derived FWHM values for the two He isotopes are quite different. This result is at variance with the experimental observation [6] that, for ternary particles of the same charge, the widths of the energy distributions are independent of mass. Hence, this difference in the FWHM values is not reasonable. Furthermore, when method 1 is applied to the data of Loveland for ²⁵²Cf [7] and of D'hondt *et al.* and Caitucoli *et al.* for $^{235}U(n,f)$ [10,11], where the α spectra extend to very low energy (~ 0.1 MeV), the derived ternary ⁵He yields become much larger than the corresponding ⁴He yields. Also, these fits result in very different widths distributions (FWHM) for the two He isotopes. These observations lead one to question the results of Ref. [4], and in particular the reported ⁵He energy.

A second approach to the fitting procedure was then employed with the only requirement that the widths of the two Gaussian distributions be the same, hereby fulfilling the requirement derived from Ref. [6]. This approach neglects the fact that the breakup of ⁵He contributes a small, additional spread in momentum to the distribution imparted by the ternary fission process. The results given in the lower part of Table I are presented under the assumption that ⁴He ternary fission dominates the alpha particle spectra, i.e., that ⁵He is

TABLE II. Yield ratios extracted by using the method 2 for the ternary ⁴He and ⁵He particles from the α spectrum of ²⁵²Cf [7] and ²³⁵U(*n*,*f*) [6,10,11].

Nuclei	Ternary particle	$\langle E \rangle$ MeV	FWHM MeV	Yield ratio	χ^2 /DOF
²⁵² Cf	⁴ He	$16.0^{+0.3}_{-0.3}$	$9.7^{-0.3}_{+0.3}$	$78.9^{-4.1}_{+4.9}$	0.91
	⁵ He	$11.0^{+0.5}_{-1.0}$	$9.7^{-0.3}_{+0.3}$	$21.1^{+3.9}_{-4.5}$	
²³⁵ U	⁴ He	$15.8^{+0.4}_{-0.3}$	$9.1^{-0.3}_{+0.4}$	$87.7^{-5.4}_{+4.7}$	0.56
	⁵ He	$11.6^{+0.9}_{-1.4}$	$9.1^{-0.3}_{+0.4}$	$12.3^{+5.2}_{-4.5}$	

associated with the component with the smaller intensity. The results of the fit are also displayed in the top part of Fig. 3. Under these fitting conditions, the results differ markedly from those obtained with method 1. The energies associated with the two particles are now quite different, even though the centroid of the low-energy Gaussian is determined with less than desirable accuracy. This is due to the low-energy cutoff of the present α spectra. This cutoff also impacts severely the accuracy with which the intensity of the two components can be determined. In other words, the low-energy cutoff of the present α spectra makes the errors large in the yields and energies as shown in Table I. These findings prompted us to then concentrate on the data of Loveland [7] and the results of a fit (with the same constraints) of the spectrum measured by this author are presented in Table II and the middle part of Fig. 3. Within the errors, the results of the two data sets are in satisfactory agreement. Furthermore, the effect of the low-energy cutoff in the new data was examined by arbitrarily truncating the data by Loveland [6]. It was found that the changes in the values of the fitting parameters remained within errors for cutoff energies varying from ~ 1 to 9.5 MeV; i.e., the values presented in Table II are quite stable and reliable. Thus, this analysis indicates that the average energies of the alpha particles associated with ⁴He and ⁵He differ by 3-6 MeV and that at least 80% of the total alpha yield is associated with ⁴He originating from ternary



FIG. 3. Two Gaussian fits with the only condition of the same FWHMs for two Gaussians. The spectrum (a) is from the present work. (b) and (c) are the alpha spectra from data of Loveland for 252 Cf [7] and from data for 235 U(*n*,*f*) of D'hondt *et al.* [10] and Caitucoli *et al.* [11], respectively.

fission. Finally, a similar analysis of the ${}^{235}U(n,f)$ ternary fission data of Refs. [10] and [11] was also carried out. The results form the last entries to Table II. As can be seen the results are similar to those obtained for ²⁵²Cf and a satisfactory understanding of the data appears to emerge. The agreement between the various data sets can be regarded as satisfactory, considering that the comparisons cover data obtained with independent instruments under different experimental conditions. The present yield of ternary ⁵He for 235 U(*n*,*f*) is $12.3^{+5.2}_{-4.5}$ % of that of ternary ⁴He, which can be compared with the $\sim 5.0\%$ predicted in the yield versus E_c plot by Halpern [5]. The present yield of $12.3^{+5.2}_{-4.5}$ % is still within the range of the yield values which are expected from a simple statistical model of Halpern [5]. Considering the uncertainty in the calculated E_c values and the errors in the measured yields of the ternary particles (see Fig. 5 in Ref. [5]) the new results are consistent with that of Halpern's prediction. The previous ⁴He yields are the summed yield of ⁵He and ⁴He. Relative yields for other ternary particles have to be increased a little, when this correction is taken into account.

To summarize, the direct yields of ⁴He and ⁵He particles

- L. W. Alvarez, as reported by G. Farwell, E. Segre, and C. Wiegand, Phys. Rev. 71, 327 (1947).
- [2] M. Mutterer, P. Singer, Yu. Kopach, M. Klemens, A. Hotzel, D. Schwalm, P. Thirolf, and M. Hesse, in *Proceedings of Dy*namical Aspects of Nuclear Fission, Casta-Papiernicka, Slovak Republic, 1996 (JINR, Dubna, 1996), p. 250.
- [3] S. W. Cosper, J. Cerny, and R. C. Gatti, Phys. Rev. 154, 1193 (1967).
- [4] E. Cheifetz, B. Eylon, E. Fraenkel, and A. Gavron, Phys. Rev. Lett. 29, 805 (1972).
- [5] I. Halpern, Annu. Rev. Nucl. Sci. 21, 245 (1971).
- [6] C. Wagemans, The Nuclear Fission Process (CRC Press, Lon-

from ternary spontaneous fission (SF) of 252 Cf have been obtained from a data set measured recently at Gammasphere. While the total α energy distribution cannot be fitted well with a single Gaussian distribution, it can be reproduced satisfactorily by assuming the presence of two components in the spectrum associated with ⁴He and ⁵He ternary fission, respectively. A consistent description of the present 252 Cf data as well as of the data of Ref. [7] for 252 Cf and of Refs. [10,11] for 235 U(*n*,*f*) was achieved. In all cases the energy of ⁵He is determined to be around 11 MeV while the corresponding energy for ⁴He is 16 MeV. 80–90% of the total yield can be assigned the ternary fission accompanied by a ⁴He particle. The remaining 10–20% of the alpha yield is then assigned the breakup of ⁵He.

Research at Vanderbilt University and Mississippi State University is supported in part by the U.S. Department of Energy under Grants No. DE-FG05-88ER40407 and DE-FG05-95ER40939. Work at Idaho National Engineering Laboratory is supported by the U.S. Department of Energy under Contract No. DE-AC07- 76ID01570. Work at Argonne National Laboratory is supported by the Department of Energy under Contract No. W-31-109-ENG-38.

don, 1991).

- [7] W. Loveland, Phys. Rev. C 9, 395 (1974).
- [8] Hongyin Han, Shengnian Huang, Jiangchen Meng, Zeongyn Bao, and Zougyuan Ye, *IAEA Consults Meet. Physics of Neu*tron Emission in Fission, edited by H. D. Lemmel (IAEA, Vienna, 1989).
- [9] I. Y. Lee, Nucl. Phys. A520, 641c (1990).
- [10] P. D'hondt, A. De Clercq, A. Deruytter, C. Wagemans, M. Asghar, and A. Emsallem, Nucl. Phys. A303, 275 (1978).
- [11] F. Caitucoli, B. Leroux, G. Barreau, N. Carjan, T. Benfoughal, T. Doan, F. El Hage, A. Sicre, M. Asghar, P. Perrin, and G. Siegert, Z. Phys. A **298**, 219 (1980).